

# PRINCIPLES & PRACTICE OF PHYSICS

Eric Mazur

ALWAYS LEARNING



## PRINCIPLES & PRACTICE OF PHYSICS BICS GLOBAL EDITION

Eric Mazur Harvard University

With contributions from

Catherine H. Crouch Swarthmore College

Peter A. Dourmashkin Massachusetts Institute of Technology

Global Edition contributions from

Andrea Banfi University of Sussex

Najeh M. Jisrawi University of Sharjah

Girish S. Setlur Indian Institute of Technology Guwahati



Boston Columbus Indianapolis New York San Francisco Hoboken Amsterdam Cape Town Dubai London Madrid Milan Munich Paris Montréal Toronto Delhi Mexico City São Paulo Sydney Hong Kong Seoul Singapore Taipei Tokyo

Executive Editor: Becky Ruden Text Permissions Specialist: Paul Sarkis Publisher: Jim Smith Associate Content Producer: Megan Power Project Managers: Beth Collins and Martha Steele Production Management: Rose Kernan Program Manager: Katie Conley Copyeditor: Carol Reitz Managing Development Editor: Cathy Murphy Compositor: Cenveo® Publisher Services Senior Development Editor: Margot Otway Design Manager: Mark Ong Development Editor: Irene Nunes Interior Designer: Hespenheide Design Editorial Assistant: Sarah Kaubisch Cover Designer: Lumina Datamatics Ltd Head of Learning Asset Acquisition, Global Edition: Laura Dent Illustrators: Rolin Graphics Senior Acquisitions Editor, Global Edition: Privanka Ahuja Photo Permissions Management: Mava Melenchuk Project Editors, Global Edition: KK Neelakantan and Photo Researcher: Eric Schrader Mrithyunjayan Nilayamgode Manufacturing Buyer: Jeff Sargent Media Producer, Global Edition: M Vikram Kumar Vice-President of Marketing: Christy Lesko Senior Manufacturing Controller, Global Edition: Kay Holman Marketing Manager: Will Moore Text Permissions Project Manager: Liz Kincaid Senior Marketing Development Manager: Michelle Cadden

Cover Photo Credit: Sheff/Shutterstock

Credits and acknowledgments borrowed from other sources and reproduced, with permission, in this textbook appear on the appropriate page within the text or on p. 1045.

Pearson Education Limited Edinburgh Gate Harlow Essex CM20 2JE England

and Associated Companies throughout the world

Visit us on the World Wide Web at: www.pearsonglobaleditions.com

© Pearson Education Limited 2015

The rights of Eric S. Mazur to be identified as the author of this work has been asserted by him in accordance with the Copyright, Designs and Patents Act 1988.

Authorized adaptation from the United States edition, entitled Principles & Practice of Physics, ISBN 978-0-321-94920-2, by Eric Mazur, published by Pearson Education © 2015.

All rights reserved. No part of this publication may be reproduced, stored in a retrieval system, or transmitted in any form or by any means, electronic, mechanical, photocopying, recording or otherwise, without either the prior written permission of the publisher or a license permitting restricted copying in the United Kingdom issued by the Copyright Licensing Agency Ltd, Saffron House, 6–10 Kirby Street, London EC1N 8TS.

Many of the designations used by manufacturers and sellers to distinguish their products are claimed as trademarks. Where those designations appear in this book, and the publisher was aware of a trademark claim, the designations have been printed in initial caps or all caps.

MasteringPhysics is a trademark, in the U.S. and/or other countries, of Pearson Education, Inc. or its affiliates.

ISBN 10: 1-292-07886-3 ISBN 13: 978-1-292-07886-1 (Print) ISBN 13: 978-1-292-07644-7 (PDF)

British Library Cataloguing-in-Publication Data A catalogue record for this book is available from the British Library

 $10\ 9\ 8\ 7\ 6\ 5\ 4\ 3\ 2\ 1$ 

Typeset in Times Ten by Cenveo® Publisher Services

Printed and bound in China

### **Brief Contents**

Chapter 1	Foundations 19			
Chapter 2	Motion in One Dimension 46			
Chapter 3	Acceleration 71			
Chapter 4	Acceleration 71 Momentum 93			
Chapter 5	Energy 119			
Chapter 6	Principle of Relativity 139			
Chapter 7	Interactions 166			
Chapter 8	Force 194			
Chapter 9	Work 220			
Chapter 10	Motion in a Plane 244			
Chapter 11	Motion in a Circle 272			
Chapter 12	Torque 299			
Chapter 13	Gravity 326			
Chapter 14	Special Relativity 355			
Chapter 15	Periodic Motion 392			
Chapter 16	Waves in One Dimension 418			
Chapter 17	-			
Chapter 18	Fluids 481			
Chapter 19	Entropy 519			
Chapter 20	Energy Transferred Thermally 547			
Chapter $21$	Degradation of Energy 580			
Chapter 22	Electric Interactions 611			
Chapter 23	The Electric Field 633			
Chapter $24$	Gauss's Law 657			
Chapter 25	Work and Energy in Electrostatics 681			
Chapter 26	Charge Separation and Storage 703			
Chapter 27	Magnetic Interactions 728			
Chapter 28	Magnetic Fields of Charged Particles in Motion 753			
Chapter 29	Changing Magnetic Fields 777			
Chapter 30	Changing Electric Fields 799			
Chapter 31	Electric Circuits 829			
Chapter 32	Electronics 860			
Chapter 33	Ray Optics 893			
Chapter $34$	Wave and Particle Optics 926			

### **About the Author**



Fric Mazur is the Balkanski Professor of Physics and Applied Physics at Harvard University and Area Dean of Applied Physics. Dr. Mazur is a renowned scientist and researcher in optical physics and in education research, and a sought-after author and speaker.

Dr. Mazur joined the faculty at Harvard shortly after obtaining his Ph.D. at the University of Leiden in the Netherlands. In 2012 he was awarded an Honorary Doctorate from the École Polytechnique and the University of Montreal. He is a Member of the Royal Academy of Sciences of the Netherlands and holds honorary professorships at the Institute of Semiconductor Physics of the Chinese Academy of Sciences in Beijing, the Institute of Laser Engineering at the Beijing University of Technology, and the Beijing Normal University.

Dr. Mazur has held appointments as Visiting Professor or Distinguished Lecturer at Carnegie Mellon University, the Ohio State University, the Pennsylvania State University, Princeton University, Vanderbilt University, Hong Kong University, the University of Leuven in Belgium, and National Taiwan University in Taiwan, among others.

In addition to his work in optical physics, Dr. Mazur is interested in education, science policy, outreach, and the public perception of science. In 1990 he began developing peer instruction, a method for teaching large lecture classes interactively. This teaching method has developed a large following, both nationally and internationally, and has been adopted across many science disciplines.

Dr. Mazur is author or co-author of over 250 scientific publications and holds two dozen patents. He has also written on education and is the author of *Peer Instruction: A User's Manual* (Pearson, 1997), a book that explains how to teach large lecture classes interactively. In 2006 he helped produce the award-winning DVD *Interactive Teaching*. He is the co-founder of Learning Catalytics, a platform for promoting interactive problem solving in the classroom, which is available in MasteringPhysics<sup>®</sup>.

### To the Student

Let me tell you a bit about myself.

I always knew exactly what I wanted to do. It just never worked out that way.

When I was seven years old, my grandfather gave me a book about astronomy. Growing up in the Netherlands I became fascinated by the structure of the solar system, the Milky Way, the universe. I remember struggling with the concept of infinite space and asking endless questions without getting satisfactory answers. I developed an early passion for space and space exploration. I knew I was going to be an astronomer. In high school I was good at physics, but when I entered university and had to choose a major, I chose astronomy.

It took only a few months for my romance with the heavens to unravel. Instead of teaching me about the mysteries and structure of the universe, astronomy had been reduced to a mind-numbing web of facts, from declinations and right ascensions to semi-major axes and eccentricities. Disillusioned about astronomy, I switched majors to physics. Physics initially turned out to be no better than astronomy, and I struggled to remain engaged. I managed to make it through my courses, often by rote memorization, but the beauty of science eluded me.

It wasn't until doing research in graduate school that I rediscovered the beauty of science. I knew one thing for sure, though: I was never going to be an academic. I was going to do something useful in my life. Just before obtaining my doctorate, I lined up my dream job working on the development of the compact disc, but I decided to spend one year doing postdoctoral research first.

It was a long year. After my postdoc, I accepted a junior faculty position and started teaching. That's when I discovered that the combination of doing research—uncovering the mysteries of the universe—and teaching—helping others to see the beauty of the universe—is a wonderful combination.

When I started teaching, I did what all teachers did at the time: lecture. It took almost a decade to discover that my award-winning lecturing did for my students exactly what the courses I took in college had done for me: It turned the subject that I was teaching into a collection of facts that my students memorized by rote. Instead of transmitting the beauty of my field, I was essentially regurgitating facts to my students.

When I discovered that my students were not mastering even the most basic principles, I decided to completely change my approach to teaching. Instead of lecturing, I asked students to read my lecture notes at home, and then, in class, I taught by questioning—by asking my students to reflect on concepts, discuss in pairs, and experience their own "aha!" moments.

Over the course of more than twenty years, the lecture notes have evolved into this book. Consider this book to be my best possible "lecturing" to you. But instead of listening to me without having the opportunity to reflect and think, this book will permit you to pause and think; to hopefully experience many "aha!" moments on your own.

I hope this book will help you develop the thinking skills that will make you successful in your career. And remember: your future may be—and likely will be—very different from what you imagine.

I welcome any feedback you have. Feel free to send me email or tweets.

I wrote this book for you.

Eric Mazur ♥@eric\_mazur mazur@harvard.edu Cambridge, MA

### **To the Instructor**

They say that the person who teaches is the one who learns the most in the classroom. Indeed, teaching led me to many unexpected insights. So, also, with the writing of this book, which has been a formidably exciting intellectual journey.

#### Why write a new physics text?

In May 1993 I was driving to Troy, NY, to speak at a meeting held in honor of Robert Resnick's retirement. In the car with me was a dear friend and colleague, Albert Altman, professor at the University of Massachusetts, Lowell. He asked me if I was familiar with the approach to physics taken by Ernst Mach in his popular lectures. I wasn't. Mach treats conservation of momentum before discussing the laws of motion, and his formulation of mechanics had a profound influence on Einstein.

The idea of using conservation principles derived from experimental observations as the basis for a text—rather than Newton's laws and the concept of force—appealed to me immediately. After all, most physicists never use the concept of force because it relates only to mechanics. It has no role in quantum physics, for example. The conservation principles, however, hold throughout all of physics. In that sense they are much more fundamental than Newton's laws. Furthermore, conservation principles involve only algebra, whereas Newton's second law is a differential equation.

It occurred to me, however, that Mach's approach could be taken further. Wouldn't it be nice to start with conservation of both momentum *and* energy, and only *later* bring in the concept of force? After all, physics education research has shown that the concept of force is fraught with pitfalls. What's more, after tediously deriving many results using kinematics and dynamics, most physics textbooks show that you can derive the same results from conservation principles in just one or two lines. Why not do it the easy way first?

It took me many years to reorganize introductory physics around the conservation principles, but the resulting approach is one that is much more unified and modern—the conservation principles are the theme that runs throughout this entire book.

Additional motives for writing this text came from my own teaching. Most textbooks focus on the acquisition of information and on the development of procedural knowledge. This focus comes at the expense of conceptual understanding or the ability to transfer knowledge to a new context. As explained below, I have structured this text to redress that balance. I also have drawn deeply on the results of physics education research, including that of my own research group.

I have written this text to be accessible and easy for students to understand. My hope is that it can take on the burden of basic teaching, freeing class time for synthesis, discussion, and problem solving.

#### Setting a new standard

The tenacity of the standard approach in textbooks can be attributed to a combination of inertia and familiarity. Teaching large introductory courses is a major chore, and once a course is developed, changing it is not easy. Furthermore, the standard texts worked for *us*, so it's natural to feel that they should work for our students, too.

The fallacy in the latter line of reasoning is now wellknown thanks to education research. Very few of our students are like us at all. Most take physics because they are required to do so; many will take no physics beyond the introductory course. Physics education research makes it clear that the standard approach fails these students.

Because of pressure on physics departments to deliver better education to non-majors, changes are occurring in the way physics is taught. These changes, in turn, create a need for a textbook that embodies a new educational philosophy in both format and presentation.

#### Organization of this book

As I considered the best way to convey the conceptual framework of mechanics, it became clear that the standard curriculum truly deserved to be rethought. For example, standard texts are forced to redefine certain concepts more than once—a strategy that we know befuddles students. (Examples are *work*, the standard definition of which is incompatible with the first law of thermodynamics, and *energy*, which is redefined when modern physics is discussed.)

Another point that has always bothered me is the arbitrary division between "modern" and "classical" physics. In most texts, the first thirty-odd chapters present physics essentially as it was known at the end of the 19th century; "modern physics" gets tacked on at the end. There's no need for this separation. Our goal should be to explain physics in the way that works best for students, using our full contemporary understanding. *All* physics is modern!

That is why my table of contents departs from the "standard organization" in the following specific ways.

**Emphasis on conservation laws.** As mentioned earlier, this book introduces the conservation laws early and treats them the way they should be: as the backbone of physics. The advantages of this shift are many. First, it avoids many of the standard pitfalls related to the concept of force, and it leads naturally to the two-body character of forces and the laws of motion. Second, the conservation laws enable students to solve a wide variety of problems without any calculus. Indeed, for complex systems, the conservation laws are often the natural (or only) way to solve problems. Third, the book deduces the conservation laws from observations, helping to make clear their connection with the world around us.

Торіс	Chapters	Can be inserted after chapter	Chapters that can be omitted without affecting continuity
Mechanics	1-14		6, 13–14
Waves	15-17	12	16–17
Fluids	18	9	
Thermal Physics	19-21	10	21
Electricity & Magnetism	22-30	12 (but 17 is needed for 29–30)	29-30
Circuits	31-32	26 (but 30 is needed for 32)	32
Optics	33-34	17	34

#### Table 1 Scheduling matrix

I and several other instructors have tested this approach extensively in our classes and found markedly improved performance on problems involving momentum and energy, with large gains on assessment instruments like the Force Concept Inventory.

**Early emphasis on the concept of system.** Fundamental to most physical models is the separation of a system from its environment. This separation is so basic that physicists tend to carry it out unconsciously, and traditional texts largely gloss over it. This text introduces the concept in the context of conservation principles and uses it consistently.

**Postponement of vectors.** Most introductory physics concerns phenomena that take place along one dimension. Problems that involve more than one dimension can be broken down into one-dimensional problems using vectorial notation. So a solid understanding of physics in one dimension is of fundamental importance. However, by introducing vectors in more than one dimension from the start, standard texts distract the student from the basic concepts of kinematics.

In this book, I develop the complete framework of mechanics for motions and interactions in one dimension. I introduce the second dimension when it is needed, starting with rotational motion. Hence, students are free to concentrate on the actual physics.

**Just-in-time introduction of concepts.** Wherever possible, I introduce concepts only when they are necessary. This approach allows students to put ideas into immediate practice, leading to better assimilation.

**Integration of modern physics.** A survey of syllabi shows that less than half the calculus-based courses in the United States cover modern physics. I have therefore integrated selected "modern" topics throughout the text. For example, special relativity is covered in Chapter 14, at the end of mechanics. Chapter 32, Electronics, includes sections on semiconductors and semiconductor devices. Chapter 34, Wave and Particle Optics, contains sections on quantization and photons.

**Modularity.** I have written the book in a modular fashion so it can accommodate a variety of curricula (See Table 1, "Scheduling matrix").

The book contains two major parts, Mechanics and Electricity and Magnetism, plus five shorter parts. The two major parts by themselves can support an in-depth two-semester or three-quarter course that presents a complete picture of physics embodying the fundamental ideas of modern physics. Additional parts can be added for a longer or faster-paced course. The five shorter parts are more or less self-contained, although they do build on previous material, so their placement is flexible. Within each part or chapter, more advanced or difficult material is placed at the end.

#### Pedagogy

This text draws on many models and techniques derived from my own teaching and from physics education research. The following are major themes that I have incorporated throughout.

Separation of conceptual and mathematical frameworks. Each chapter is divided into two parts: Concepts and Quantitative Tools. The first part, Concepts, develops the full conceptual framework of the topic and addresses many of the common questions students have. It concentrates on the underlying ideas and paints the big picture, whenever possible without equations. The second part of the chapter, Quantitative Tools, then develops the mathematical framework.

**Deductive approach; focus on ideas before names and equations.** To the extent possible, this text develops arguments deductively, starting from observations, rather than stating principles and then "deriving" them. This approach makes the material easier to assimilate for students. In the same vein, this text introduces and explains each idea before giving it a formal name or mathematical definition.

Stronger connection to experiment and experience. Physics stems from observations, and this text is structured so that it can do the same. As much as possible, I develop the material from experimental observations (and preferably those that students can make) rather than assertions. Most chapters use actual data in developing ideas, and new notions are always introduced by going from the specific to the general whenever possible by interpreting everyday examples.

By contrast, standard texts often introduce laws in their most general form and then show that these laws are consistent with specific (and often highly idealized) cases. Consequently the world of physics and the "real" world remain two different things in the minds of students. Addressing physical complications. I also strongly oppose presenting unnatural situations; real life complications must always be confronted head-on. For example, the use of unphysical words like *frictionless* or *massless* sends a message to the students that physics is unrealistic or, worse, that the world of physics and the *real* world are unrelated entities. This can easily be avoided by pointing out that friction or mass may be neglected under certain circumstances and pointing out *why* this may be done.

**Engaging the student.** Education is more than just transfer of information. Engaging the student's mind so the information can be assimilated is essential. To this end, the text is written as a dialog between author and reader (often invoking the reader—*you*—in examples) and is punctuated by Checkpoints—questions that require the reader to stop and think. The text following a Checkpoint often refers directly to its conclusions. Students will find complete solutions to all the Checkpoints at the back of the book; these solutions are written to emphasize physical reasoning and discovery.

**Visualization.** Visual representations are central to physics, so I developed each chapter by designing the figures before writing the text. Many figures use multiple representations to help students make connections (for example, a sketch may be combined with a graph and a bar diagram). Also, in accordance with research, the illustration style is spare and simple, putting the emphasis on the ideas and relationships rather than on irrelevant details. The figures do not use perspective unless it is needed, for instance.

#### Structure of this text

#### Division into Principles and Practice books

I've divided this text into a *Principles* book, which teaches the physics, and a *Practice* book, which puts the physics into practice and develops problem-solving skills. This division helps address two separate intellectually demanding tasks: understanding the physics and learning to solve problems. When these two tasks are mixed together, as they are in standard texts, students are easily overwhelmed. Consequently many students focus disproportionately on worked examples and procedural knowledge, at the expense of the physics.

#### Structure of Principles chapters

As pointed out earlier, each *Principles* chapter is divided into two parts. The first part (Concepts) develops the conceptual framework in an accessible way, relying primarily on qualitative descriptions and illustrations. In addition to including Checkpoints, each Concepts section ends with a one-page Self-quiz consisting of qualitative questions.

The second part of each chapter (Quantitative Tools) formalizes the ideas developed in the first part in mathematical terms. While concise, it is relatively traditional in nature teachers should be able to continue to use material developed for earlier courses. To avoid creating the impression that equations are more important than the concepts behind them, no equations are highlighted or boxed.

Both parts of the *Principles* chapters contain worked examples to help students develop problem-solving skills.

#### Structure of the Practice chapters

This book contains material to put into practice the concepts and principles developed in the corresponding chapters in the *Principles* book. Each chapter contains the following sections:

- 1. *Chapter Summary*. This section provides a brief tabular summary of the material presented in the corresponding *Principles* chapter.
- 2. *Review Questions.* The goal of this section is to allow students to quickly review the corresponding *Principles* chapter. The questions are straightforward one-liners starting with "what" and "how" (rather than "why" or "what if").
- 3. Developing a Feel. The goals of this section are to develop a quantitative feel for the quantities introduced in the chapter; to connect the subject of the chapter to the real world; to train students in making estimates and assumptions; to bolster students' confidence in dealing with unfamiliar material. It can be used for self-study or for a homework or recitation assignment. This section, which has no equivalent in existing books, combines a number of ideas (specifically, Fermi problems and tutoring in the style of the Princeton Learning Guide). The idea is to start with simple estimation problems and then build up to Fermi problems (in early chapters Fermi problems are hard to compose because few concepts have been introduced). Because students initially find these questions hard, the section provides many hints, which take the form of questions. A key then provides answers to these "hints."
- **4.** *Worked and Guided Problems.* This section contains complex worked examples whose primary goal is to teach problem solving. The Worked Problems are fully solved; the Guided Problems have a list of questions and suggestions to help the student think about how to solve the problem. Typically, each Worked Problem is followed by a related Guided Problem.
- 5. Questions and Problems. This is the chapter's problem set. The problems 1) offer a range of levels, 2) include problems relating to client disciplines (life sciences, engineering, chemistry, astronomy, etc.), 3) use the second person as much as possible to draw in the student, and 4) do not spoon-feed the students with information and unnecessary diagrams. The problems are classified into three levels as follows: (•) application of single concept; numerical plugand-chug; (••) nonobvious application of single concept or application of multiple concepts from current chapter; straightforward numerical or algebraic computation; (••) application of multiple concepts, possibly spanning multiple chapters. Context-rich problems are designated CR.

As I was developing and class-testing this book, my students provided extensive feedback. I have endeavored to

incorporate all of their feedback to make the book as useful as possible for future generations of students. In addition, the book was class-tested at a large number of institutions, and many of these institutions have reported significant increases in learning gains after switching to this manuscript. I am confident the book will help increase the learning gains in your class as well. It will help you, as the instructor, coach your students to be the best they can be.

#### Instructor supplements

The Instructor Resource Collection available via the Instructor Resource Center includes an Image Library, the Procedure and special topic boxes from *Principles*, and a library of presentation applets from ActivPhysics, PhET simulations, and PhET Clicker Questions. Lecture Outlines with embedded Clicker Questions in PowerPoint<sup>®</sup> are provided, as well as the *Instructor's Guide* and *Instructor's Solutions Manual*.

The *Instructor's Guide* provides chapter-by-chapter ideas for lesson planning using *Principles & Practice of Physics* in class, including strategies for addressing common student difficulties.

The *Instructor's Solutions Manual* is a comprehensive solutions manual containing complete answers and solutions to all Developing a Feel questions, Guided Problems, and Questions and Problems from the *Practice* book. The solutions to the Guided Problems use the book's four-step problem-solving strategy (Getting Started, Devise Plan, Execute Plan, Evaluate Result).

MasteringPhysics<sup>®</sup> is the leading online homework, tutorial, and assessment product designed to improve results by helping students quickly master concepts. Students benefit from self-paced tutorials that feature specific wrong-answer feedback, hints, and a wide variety of educationally effective content to keep them engaged and on track. Robust diagnostics and unrivalled gradebook reporting allow instructors to pinpoint the weaknesses and misconceptions of a student or class to provide timely intervention.

MasteringPhysics enables instructors to:

- Easily assign **tutorials** that provide individualized coaching.
- Mastering's hallmark **Hints** and **Feedback** offer scaffolded instruction similar to what students would experience in an office hour.

- **Hints** (declarative and Socratic) can provide problemsolving strategies or break the main problem into simpler exercises.
- Feedback lets the student know precisely what misconception or misunderstanding is evident from their answer and offers ideas to consider when attempting the problem again.

Learning Catalytics<sup>™</sup> is a "bring your own device" student engagement, assessment, and classroom intelligence system available within MasteringPhysics. With Learning Catalytics you can:

- Assess students in real time, using open-ended tasks to probe student understanding.
- Understand immediately where students are and adjust your lecture accordingly.
- Improve your students' critical-thinking skills.
- Access rich analytics to understand student performance.
- Add your own questions to make Learning Catalytics fit your course exactly.
- Manage student interactions with intelligent grouping and timing.

The **Test Bank** contains more than 2000 high-quality problems, with a range of multiple-choice, true-false, shortanswer, and conceptual questions correlated to *Principles & Practice of Physics* chapters. Test files are provided in both TestGen<sup>®</sup> and Microsoft<sup>®</sup> Word for Mac and PC.

Instructor supplements are available on the Instructor Resource Center at www.pearsonglobaleditions.com/mazur, and in the Instructor Resource area at www.masteringphysics .com.

#### Student supplements

**MasteringPhysics** (www.masteringphysics.com) is designed to provide students with customized coaching and individualized feedback to help improve problem-solving skills. Students complete homework efficiently and effectively with tutorials that provide targeted help.

**Interactive eText** allows you to highlight text, add your own study notes, and review your instructor's personalized notes, 24/7. The eText is available through MasteringPhysics, www.masteringphysics.com.

### **Acknowledgments**

This book would not exist without the contributions from many people. It was Tim Bozik, currently President, Higher Education at Pearson plc, who first approached me about writing a physics textbook. If it wasn't for his persuasion and his belief in me, I don't think I would have ever undertaken the writing of a textbook. Tim's suggestion to develop the art electronically also had a major impact on my approach to the development of the visual part of this book.

Albert Altman pointed out Ernst Mach's approach to developing mechanics starting with the law of conservation of momentum. Al encouraged me throughout the years as I struggled to reorganize the material around the conservation principles.

I am thankful to Irene Nunes, who served as Development Editor through several iterations of the manuscript. Irene forced me to continuously rethink what I had written and her insights in physics kept surprising me. Her incessant questioning taught me that one doesn't need to be a science major to obtain a deep understanding of how the world around us works and that it is possible to explain physics in a way that makes sense for non-physics majors.

Catherine Crouch helped write the final chapters of electricity and magnetism and the chapters on circuits and optics, permitting me to focus on the overall approach and the art program. Peter Dourmashkin helped me write the chapters on special relativity and thermodynamics. Without his help, I would not have been able to rethink how to introduce the ideas of modern physics in a consistent way.

Many people provided feedback during the development of the manuscript. I am particularly indebted to the late Ronald Newburgh and to Edward Ginsberg, who meticulously checked many of the chapters. I am also grateful to Edwin Taylor for his critical feedback on the special relativity chapter and to my colleague Gary Feldman for his suggestions for improving that chapter.

Lisa Morris provided material for many of the Self-quizzes and my graduate students James Carey, Mark Winkler, and Ben Franta helped with data analysis and the appendices. I would also like to thank my uncle, Erich Lessing, for letting me use some of his beautiful pictures as chapter openers.

Many people helped put together the *Practice* book. Without Daryl Pedigo's hard work authoring and editing content, as well as coordinating the contributions to that book, the manuscript would never have taken shape. Along with Daryl, the following people provided the material for the *Practice* book: Wayne Anderson, Bill Ashmanskas, Linda Barton, Ronald Bieniek, Michael Boss, Anthony Buffa, Catherine Crouch, Peter Dourmashkin, Paul Draper, Andrew Duffy, Edward Ginsberg, William Hogan, Gerd Kortemeyer, Rafael Lopez-Mobilia, Christopher Porter, David Rosengrant, Gay Stewart, Christopher Watts, Lawrence Weinstein, Fred Wietfeldt, and Michael Wofsey.

I would also like to thank the editorial and production staff at Pearson. Margot Otway helped realize my vision for the art program. Martha Steele and Beth Collins made sure the production stayed on track. In addition, I would like to thank Frank Chmely for his meticulous accuracy checking of the manuscript. I am indebted to Jim Smith and Becky Ruden for supporting me through the final stages of this process and to Carol Trueheart, Alison Reeves, and Christian Botting of Prentice Hall for keeping me on track during the early stages of the writing of this book. Finally, I am grateful to Will Moore for his enthusiasm in developing the marketing program for this book.

I am also grateful to the participants of the NSF Faculty Development Conference "Teaching Physics Conservation Laws First" held in Cambridge, MA, in 1997. This conference helped validate and cement the approach in this book.

Finally, I am indebted to the hundreds of students in Physics 1, Physics 11, and Applied Physics 50 who used early versions of this text in their course and provided the feedback that ended up turning my manuscript into a text that works not just for instructors but, more importantly, for students.

### Reviewers of *Principles & Practice of Physics*

Over the years many people reviewed and class-tested the manuscript. The author and publisher are grateful for all of the feedback the reviewers provided, and we apologize if there are any names on this list that have been inadvertently omitted.

Edward Adelson, Ohio State University Albert Altman, University of Massachusetts, Lowell Susan Amador Kane, Haverford College James Andrews, Youngstown State University Arnold Arons, University of Washington Robert Beichner, North Carolina State University Bruce Birkett, University of California, Berkeley David Branning, Trinity College Bernard Chasan, Boston University Stéphane Coutu, Pennsylvania State University Corbin Covault, Case Western Reserve University Catherine Crouch, Swarthmore College Paul D'Alessandris, Monroe Community College Paul Debevec, University of Illinois at Urbana-Champaign N. John DiNardo, Drexel University Margaret Dobrowolska-Furdyna, Notre Dame University Paul Draper, University of Texas, Arlington David Elmore, Purdue University Robert Endorf, University of Cincinnati Thomas Furtak, Colorado School of Mines Ian Gatland, Georgia Institute of Technology J. David Gavenda, University of Texas, Austin Edward Ginsberg, University of Massachusetts, Boston Gary Gladding, University of Illinois Christopher Gould, University of Southern California Victoria Greene, Vanderbilt University Benjamin Grinstein, University of California, San Diego Kenneth Hardy, Florida International University Gregory Hassold, Kettering University Peter Heller, Brandeis University Laurent Hodges, Iowa State University Mark Holtz, Texas Tech University Zafar Ismail, Daemen College Ramanathan Jambunathan, University of Wisconsin Oshkosh Brad Johnson, Western Washington University Dorina Kosztin, University of Missouri Columbia Arthur Kovacs, Rochester Institute of Technology (deceased) Dale Long, Virginia Polytechnic Institute (deceased)

#### John Lyon, Dartmouth College Trecia Markes, University of Nebraska, Kearney Peter Markowitz, Florida International University Bruce Mason, University of Oklahoma John McCullen, University of Arizona James McGuire, Tulane University Timothy McKay, University of Michigan Carl Michal, University of British Columbia Kimball Milton, University of Oklahoma Charles Misner, University of Maryland, College Park Sudipa Mitra-Kirtley, Rose-Hulman Institute of Technology Delo Mook, Dartmouth College Lisa Morris, Washington State University Edmund Myers, Florida State University Alan Nathan, University of Illinois K.W. Nicholson, Central Alabama Community College Fredrick Olness, Southern Methodist University Dugan O'Neil, Simon Fraser University Patrick Papin, San Diego State University George Parker, North Carolina State University Claude Penchina, University of Massachusetts, Amherst William Pollard, Valdosta State University Amy Pope, Clemson University Joseph Priest, Miami University (deceased) Joel Primack, University of California, Santa Cruz Rex Ramsier, University of Akron Steven Rauseo, University of Pittsburgh Lawrence Rees, Brigham Young University Carl Rotter, West Virginia University Leonard Scarfone, University of Vermont Michael Schatz, Georgia Institute of Technology Cindy Schwarz, Vassar College Hugh Scott, Illinois Institute of Technology Janet Segar, Creighton University Shahid Shaheen, Florida State University David Sokoloff, University of Oregon Gay Stewart, University of Arkansas Roger Stockbauer, Louisiana State University William Sturrus, Youngstown State University Carl Tomizuka, University of Illinois Mani Tripathi, University of California-Davis Rebecca Trousil, Skidmore College Christopher Watts, Auburn University Robert Weidman, Michigan Technological University Ranjith Wijesinghe, Ball State University Augden Windelborn, Northern Illinois University

#### **Reviewers of the Global Edition**

Andre E. Botha, *University of South America* D. K. Bhattacharya Sushil Kumar, *University of Delhi*  Samrat Mukherjee, Birla Institute of Technology, Mesra Stefan Nikolov, University of Plovdiv Kamil Syed, Shiv Nadar University Olav Syljuåsen, University of Oslo Hongzhou Zhang, Trinity College

### **Detailed Contents**

About the Author 4 To the Student 5 To the Instructor 6 Acknowledgments 10

#### Chapter 1 Foundations 19

- 1.1 The scientific method 20
- 1.2 Symmetry 22
- 1.3 Matter and the universe 24
- 1.4 Time and change 26
- 1.5 Representations 27
- 1.6 Physical quantities and units 32
- 1.7 Significant digits 35
- 1.8 Solving problems 38
- 1.9 Developing a feel 41

#### Chapter 2 Motion in One Dimension 46

- 2.1 From reality to model 47
- 2.2 Position and displacement 48
- 2.3 Representing motion 50
- 2.4 Average speed and average velocity 52
- 2.5 Scalars and vectors 57
- 2.6 Position and displacement vectors 59
- 2.7 Velocity as a vector 63
- 2.8 Motion at constant velocity 64
- 2.9 Instantaneous velocity 66



#### Chapter 3 Acceleration 71

- 3.1 Changes in velocity 72
- 3.2 Acceleration due to gravity 73
- 3.3 Projectile motion 75
- 3.4 Motion diagrams 77
- 3.5 Motion with constant acceleration 81

- 3.6 Free-fall equations 84
- 3.7 Inclined planes 87
- 3.8 Instantaneous acceleration 88

#### Chapter 4 Momentum 93

- 4.1 Friction 94
- 4.2 Inertia 94
- 4.3 What determines inertia? 98
- 4.4 Systems 99
- 4.5 Inertial standard 104
- 4.6 Momentum 105
- 4.7 Isolated systems 107
- 4.8 Conservation of momentum 112

#### Chapter 5 Energy 119

- 5.1 Classification of collisions 120
- 5.2 Kinetic energy 121
- 5.3 Internal energy 123
- 5.4 Closed systems 126
- 5.5 Elastic collisions 130
- 5.6 Inelastic collisions 133
- 5.7 Conservation of energy 134
- 5.8 Explosive separations 136

#### Chapter 6 Principle of Relativity 139

- 6.1 Relativity of motion 140
- 6.2 Inertial reference frames 142
- 6.3 Principle of relativity 144
- 6.4 Zero-momentum reference frame 148
- 6.5 Galilean relativity 151
- 6.6 Center of mass 155
- 6.7 Convertible kinetic energy 160
- 6.8 Conservation laws and relativity 163

#### Chapter 7 Interactions 166

- 7.1 The effects of interactions 167
- 7.2 Potential energy 170
- 7.3 Energy dissipation 171
- 7.4 Source energy 174
- 7.5 Interaction range 176
- 7.6 Fundamental interactions 178
- 7.7 Interactions and accelerations 182
- 7.8 Nondissipative interactions 183
- 7.9 Potential energy near Earth's surface 186
- 7.10 Dissipative interactions 189

#### Chapter 8 Force 194

- 8.1 Momentum and force 195
- 8.2 The reciprocity of forces 196
- 8.3 Identifying forces 198
- 8.4 Translational equilibrium 199
- 8.5 Free-body diagrams 200
- 8.6 Springs and tension 202
- 8.7 Equation of motion 206
- 8.8 Force of gravity 209
- **8.9** Hooke's law 210
- 8.10 Impulse 212
- 8.11 Systems of two interacting objects 214
- 8.12 Systems of many interacting objects 216



#### Chapter 9 Work 220

- 9.1 Force displacement 221
- 9.2 Positive and negative work 222
- 9.3 Energy diagrams 224
- 9.4 Choice of system 226
- 9.5 Work done on a single particle 231
- 9.6 Work done on a many-particle system 234
- 9.7 Variable and distributed forces 238
- 9.8 Power 241

#### Chapter 10 Motion in a Plane 244

- 10.1 Straight is a relative term 245
- 10.2 Vectors in a plane 246
- 10.3 Decomposition of forces 249
- 10.4 Friction 252
- 10.5 Work and friction 253
- 10.6 Vector algebra 256
- **10.7** Projectile motion in two dimensions 258
- 10.8 Collisions and momentum in two dimensions 260

- 10.9 Work as the product of two vectors 261
- 10.10 Coefficients of friction 266

#### Chapter 11 Motion in a Circle 272

- 11.1 Circular motion at constant speed 273
- 11.2 Forces and circular motion 277
- 11.3 Rotational inertia 280
- **11.4** Rotational kinematics 282
- 11.5 Angular momentum 287
- 11.6 Rotational inertia of extended objects 292

#### Chapter 12 Torque 299

- 12.1 Torque and angular momentum 300
- 12.2 Free rotation 303
- 12.3 Extended free-body diagrams 304
- 12.4 The vectorial nature of rotation 306
- 12.5 Conservation of angular momentum 311
- 12.6 Rolling motion 315
- 12.7 Torque and energy 320
- 12.8 The vector product 322

#### Chapter 13 Gravity 326

- 13.1 Universal gravity 327
- 13.2 Gravity and angular momentum 332
- 13.3 Weight 335
- 13.4 Principle of equivalence 338
- 13.5 Gravitational constant 343
- 13.6 Gravitational potential energy 344
- 13.7 Celestial mechanics 347
- 13.8 Gravitational force exerted by a sphere 352



#### Chapter 14 Special Relativity 355

- 14.1Time measurements356
- 14.2 Simultaneous is a relative term 359
- 14.3 Space-time 363

- 14.4 Matter and energy 368
- 14.5 Time dilation 373
- 14.6 Length contraction 378
- 14.7 Conservation of momentum 382
- 14.8 Conservation of energy 386

#### Chapter 15 Periodic Motion 392

- 15.1 Periodic motion and energy 393
- 15.2 Simple harmonic motion 395
- 15.3 Fourier's theorem 397
- 15.4 Restoring forces in simple harmonic motion 399
- 15.5 Energy of a simple harmonic oscillator 403
- 15.6 Simple harmonic motion and springs 407
- 15.7 Restoring torques 411
- 15.8 Damped oscillations 414



#### Chapter 16 Waves in One Dimension 418

- **16.1** Representing waves graphically 419
- 16.2 Wave propagation 422
- 16.3 Superposition of waves 427
- 16.4 Boundary effects 429
- 16.5 Wave functions 434
- 16.6 Standing waves 439
- 16.7 Wave speed 442
- 16.8 Energy transport in waves 444
- 16.9 The wave equation 447

#### Chapter 17 Waves in Two and Three Dimensions 451

- 17.1 Wavefronts 452
- 17.2 Sound 454
- 17.3 Interference 457
- 17.4 Diffraction 463
- 17.5 Intensity 466

- 17.6 Beats 469
- 17.7 Doppler effect 472
- 17.8 Shock waves 477



#### Chapter 18 Fluids 481

- 18.1 Forces in a fluid 482
- 18.2 Buoyancy 487
- 18.3 Fluid flow 489
- 18.4 Surface effects 493
- 18.5 Pressure and gravity 500
- 18.6 Working with pressure 505
- 18.7 Bernoulli's equation 509
- 18.8 Viscosity and surface tension 512

#### Chapter 19 Entropy 519

- 19.1 States 520
- 19.2 Equipartition of energy 523
- **19.3** Equipartition of space 525
- 19.4 Evolution toward the most probable macrostate 527
- 19.5 Dependence of entropy on volume 532
- 19.6 Dependence of entropy on energy 537
- 19.7 Properties of a monatomic ideal gas 541
- 19.8 Entropy of a monatomic ideal gas 544

#### Chapter 20 Energy Transferred Thermally 547

- 20.1 Thermal interactions 548
- 20.2 Temperature measurement 552
  - 20.3 Heat capacity 555
  - 20.4 PV diagrams and processes 559
  - 20.5 Change in energy and work 565

- 20.6 Isochoric and isentropic ideal gas processes 567
- 20.7 Isobaric and isothermal ideal gas processes 569
- 20.8 Entropy change in ideal gas processes 573
- 20.9 Entropy change in nonideal gas processes 577



#### Chapter 21 Degradation of Energy 580

- 21.1 Converting energy 581
- 21.2 Quality of energy 584
- 21.3 Heat engines and heat pumps 588
- 21.4 Thermodynamic cycles 593
- 21.5 Entropy constraints on energy transfers 598
- 21.6 Heat engine performance 601
- 21.7 Carnot cycle 605
- 21.8 Brayton cycle 607

#### Chapter 22 Electric Interactions 611

- 22.1 Static electricity 612
- 22.2 Electrical charge 613
- 22.3 Mobility of charge carriers 616
- 22.4 Charge polarization 620
- 22.5 Coulomb's law 624
- 22.6 Force exerted by distributions of charge carriers 628

#### Chapter 23 The Electric Field 633

- 23.1 The field model 634
- 23.2 Electric field diagrams 636
- 23.3 Superposition of electric fields 637
- 23.4 Electric fields and forces 640
- 23.5 Electric field of a charged particle 644
- 23.6 Dipole field 645

- 23.7 Electric fields of continuous charge distributions 647
- 23.8 Dipoles in electric fields 652

#### Chapter 24 Gauss's Law 657

- 24.1 Electric field lines 658
- 24.2 Field line density 660
- 24.3 Closed surfaces 661
- 24.4 Symmetry and Gaussian surfaces 663
- 24.5 Charged conducting objects 666
- 24.6 Electric flux 670
- 24.7 Deriving Gauss's law 672
- 24.8 Applying Gauss's law 674



#### Chapter 25 Work and Energy in Electrostatics 681

- 25.1 Electrical potential energy 682
- 25.2 Electrostatic work 683
- 25.3 Equipotentials 685
- 25.4 Calculating work and energy in electrostatics 689
- 25.5 Potential difference 692
- 25.6 Electric potentials of continuous charge distributions 698
- 25.7 Obtaining the electric field from the potential 700



#### Chapter 26 Charge Separation and Storage 703

- 26.1 Charge separation 704
- 26.2 Capacitors 706
- 26.3 Dielectrics 710
- 26.4 Voltaic cells and batteries 712
- 26.5 Capacitance 715
- 26.6 Electric field energy and emf 718
- 26.7 Dielectric constant 721
- 26.8 Gauss's law in dielectrics 725

#### Chapter 27 Magnetic Interactions 728

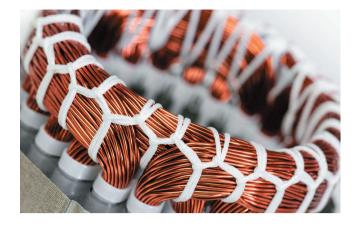
- 27.1 Magnetism 729
- 27.2 Magnetic fields 731
- 27.3 Charge flow and magnetism 733
- 27.4 Magnetism and relativity 736
- 27.5 Current and magnetism 741
- 27.6 Magnetic flux 743
- 27.7 Moving particles in electric and magnetic fields 745
- 27.8 Magnetism and electricity unified 749

#### Chapter 28 Magnetic Fields of Charged Particles in Motion 753

- 28.1 Source of the magnetic field 754
- 28.2 Current loops and spin magnetism 755
- 28.3 Magnetic dipole moment and torgue 757
- 28.4 Ampèrian paths 760
- 28.5 Ampère's law 764
- 28.6 Solenoids and toroids 767
- 28.7 Magnetic fields due to currents 770
- 28.8 Magnetic field of a moving charged particle 773

#### Chapter 29 Changing Magnetic Fields 777

- 29.1 Moving conductors in magnetic fields 778
- 29.2 Faraday's law 780
- 29.3 Electric fields accompany changing magnetic fields 781
- 29.4 Lenz's law 782
- 29.5 Induced emf 787
- 29.6 Electric field accompanying a changing magnetic field 791
- 29.7 Inductance 793
- 29.8 Magnetic energy 795

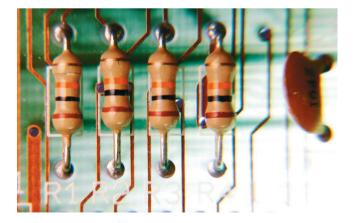


#### Chapter 30 Changing Electric Fields 799

- 30.1 Magnetic fields accompany changing electric fields 800
- **30.2** Fields of moving charged particles 803
- **30.3** Oscillating dipoles and antennas 806
- **30.4** Displacement current 812
- 30.5 Maxwell's equations 816
- 30.6 Electromagnetic waves 819
- 30.7 Electromagnetic energy 824

#### Chapter 31 Electric Circuits 829

- **31.1** The basic circuit 830
- 31.2 Current and resistance 832
- 31.3 Junctions and multiple loops 834
- 31.4 Electric fields in conductors 837
- **31.5** Resistance and Ohm's law 842
- 31.6 Single-loop circuits 846
- 31.7 Multiloop circuits 851
- **31.8** Power in electric circuits 856



#### Chapter 32 Electronics 860

- 32.1 Alternating currents 861
- 32.2 AC circuits 863
- 32.3 Semiconductors 867
- 32.4 Diodes, transistors, and logic gates 869
- 32.5 Reactance 875
- 32.6 RC and RLC series circuits 880
- 32.7 Resonance 887
- 32.8 Power in AC circuits 889

#### Chapter 33 Ray Optics 893

- 33.1 Rays 894
- 33.2 Absorption, transmission, and reflection 895
- **33.3** Refraction and dispersion 898
- 33.4 Forming images 902
- 33.5 Snel's law 909
- **33.6** Thin lenses and optical instruments **912**
- 33.7 Spherical mirrors 919
- 33.8 Lensmaker's formula 922

#### Chapter 34 Wave and Particle Optics 926

- **34.1** Diffraction of light 927
- **34.2** Diffraction gratings 929
- 34.3 X-ray diffraction 932
- 34.4 Matter waves 936
- 34.5 Photons 937
- 34.6 Multiple-slit interference 941
- 34.7 Thin-film interference 946
- 34.8 Diffraction at a single-slit barrier 948
- 34.9 Circular apertures and limits of resolution 949
- **34.10** Photon energy and momentum **953**

Appendix A: Notation 959 Appendix B: Mathematics Review 969 Appendix C: SI Units, Useful Data, and Unit Conversion Factors 975 Appendix D: Center of Mass of Extended Objects 979 Appendix E: Derivation of the Lorentz Transformation Equations 980 Solutions to Checkpoints 983 Credits 1045 Index 1047

This page is intentionally left blank.



- 1.1 The scientific method
- 1.2 Symmetry
- **1.3 Matter and the universe**
- 1.4 Time and change
- 1.5 Representations
- 1.6 Physical quantities and units
- 1.7 Significant digits
- 1.8 Solving problems
- **1.9 Developing a feel**

hances are you are taking this course in physics because someone told you to take it, and it may not be clear to you *why* you should be taking it. One good reason for taking a physics course is that, first and foremost, physics provides a fundamental understanding of the world. Furthermore, whether you are majoring in psychology, engineering, biology, physics, or something else, this course offers you an opportunity to sharpen your reasoning skills. Knowing physics means becoming a better problem solver (and I mean *real* problems, not textbook problems that have already been solved), and becoming a better problem solver is empowering: It allows you to step into unknown territory with more confidence. Before we embark on this exciting journey, let's map out the territory we are going to explore so that you know where we are going.

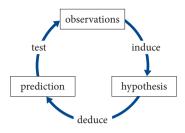
#### 1.1 The scientific method

Physics, from the Greek word for "nature," is commonly defined as the study of matter and motion. Physics is about discovering the wonderfully simple unifying patterns that underlie absolutely everything that happens around us, from the scale of subatomic particles, to the microscopic world of DNA molecules and cells, to the cosmic scale of stars, galaxies, and planets. Physics deals with atoms and molecules; gases, solids, and liquids; everyday objects, and black holes. Physics explores motion, light, and sound; the creation and annihilation of matter; evaporation and melting; electricity and magnetism. Physics is all around you: in the Sun that provides your daylight, in the structure of your bones, in your computer, in the motion of a ball you throw. In a sense, then, physics is the study of all there is in the universe. Indeed, biology, engineering, chemistry, astronomy, geology, and so many other disciplines you might name all use the principles of physics.

The many remarkable scientific accomplishments of ancient civilizations that survive to this day testify to the fact that curiosity about the world is part of human nature. Physics evolved from *natural philosophy*—a body of knowledge accumulated in ancient times in an attempt to explain the behavior of the universe through philosophical speculation—and became a distinct discipline during the scientific revolution that began in the 16th century. One of the main changes that occurred in that century was the development of the **scientific method**, an iterative process for going from observations to validated theories.

In its simplest form, the scientific method works as follows (Figure 1.1): A researcher makes a number of observations concerning either something happening in the natural world (a volcano erupting, for instance) or something happening during a laboratory experiment (a dropped brick and a dropped Styrofoam peanut travel to the floor at different speeds). These observations then lead the researcher to formulate a **hypothesis**, which is a tentative explanation of the observed phenomenon. The hypothesis is used to predict

**Figure 1.1** The scientific method is an iterative process in which a hypothesis, which is inferred from observations, is used to make a prediction, which is then tested by making new observations.



the outcome of some related natural occurrence (how a similarly shaped mountain near the erupting volcano will behave) or related laboratory experiment (what happens when a book and a sheet of paper are dropped at the same time). If the predictions prove inaccurate, the hypothesis must be modified. If the predictions prove accurate in test after test, the hypothesis is elevated to the status of either a **law** or a **theory**.

A law tells us *what* happens under certain circumstances. Laws are usually expressed in the form of relationships between observable quantities. A theory tells us *why* something happens and explains phenomena in terms of more basic processes and relationships. A scientific theory is not a mere conjecture or speculation. It is a thoroughly tested explanation of a natural phenomenon, one that is capable of making predictions that can be verified by experiment. The constant testing and retesting are what make the scientific method such a powerful tool for investigating the universe: The results obtained must be repeatable and verifiable by others.

#### Exercise 1.1 Hypothesis or not?

Which of the following statements are hypotheses? (*a*) Heavier objects fall to Earth faster than lighter ones. (*b*) The planet Mars is inhabited by invisible beings that are able to elude any type of observation. (*c*) Distant planets harbor forms of life. (*d*) Handling toads causes warts.

**SOLUTION** (*a*), (*c*), and (*d*). A hypothesis must be experimentally verifiable. (*a*) I can verify this statement by dropping a heavy object and a lighter one at the same instant and observing which one hits the ground first. (*b*) This statement asserts that the beings on Mars cannot be observed, which precludes any experimental verification and means this statement is not a valid hypothesis. (*c*) Although we humans currently have no means of exploring or closely observing distant planets, the statement is in principle testable. (*d*) Even though we know this statement is false, it *is* verifiable and therefore is a hypothesis.

Because of the constant reevaluation demanded by the scientific method, science is not a stale collection of facts but rather a living and changing body of knowledge. More important, any theory or law *always* remains tentative, and the testing never ends. In other words, it is not possible to

prove any scientific theory or law to be absolutely true (or even absolutely false). Thus the material you will learn in this book does not represent some "ultimate truth"—it is true only to the extent that it has not been proved wrong.

A case in point is *classical mechanics*, a theory developed in the 17th century to describe the motion of everyday objects (and the subject of most of this book). Although this theory produces accurate results for most everyday phenomena, from balls thrown in the air to satellites orbiting Earth, observations made during the last hundred years have revealed that under certain circumstances, significant deviations from this theory occur. It is now clear that classical mechanics is applicable for only a limited (albeit important) range of phenomena, and new branches of physics *quantum mechanics* and the theory of *special relativity* among them—are needed to describe the phenomena that fall outside the range of classical mechanics.

The formulation of a hypothesis almost always involves developing a model, which is a simplified conceptual representation of some phenomenon. You don't have to be trained as a scientist to develop models. Everyone develops mental models of how people behave, how events unfold, and how things work. Without such models, we would not be able to understand our experiences, decide what actions to take, or handle unexpected experiences. Examples of models we use in everyday life are that door handles and door hinges are on opposite sides of doors and that the + button on a TV remote increases the volume or the channel number. In everyday life, we base our models on whatever knowledge we have, real or imagined, complete or incomplete. In science we must build models based on careful observation and determine ways to fill in any missing information.

Let's look at the iterative process of developing models and hypotheses in physics, with an eye toward determining what skills are needed and what pitfalls are to be avoided (Figure 1.2). Developing a scientific hypothesis often begins

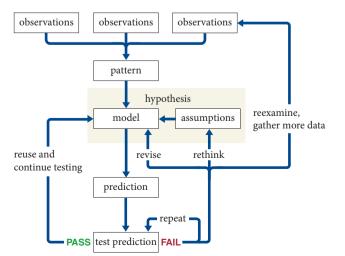


Figure 1.2 Iterative process for developing a scientific hypothesis.

with recognizing patterns in a series of observations. Sometimes these observations are direct, but sometimes we must settle for indirect observations. (We cannot directly observe the nucleus of an atom, for instance, but a physicist can describe the structure of the nucleus and its behavior with great certainty and accuracy.) As Figure 1.2 indicates, the patterns that emerge from our observations must often be combined with simplifying assumptions to build a model. The combination of model and assumptions is what constitutes a hypothesis.

It may seem like a shaky proposition to build a hypothesis on assumptions that are accepted without proof, but making these assumptions—*consciously*—is a crucial step in making sense of the universe. All that is required is that, when formulating a hypothesis, we must be aware of these assumptions and be ready to revise or drop them if the predictions of our hypothesis are not validated. We should, in particular, watch out for what are called *hidden assumptions* assumptions we make without being aware of them. As an example, try answering the following question. (Turn to the final section of the *Principles* volume, "Solutions to checkpoints," for the answer.)

**1.1** I have two coins in my pocket, together worth 30 cents. If one of them is not a nickel, what coins are they?

Advertising agencies and magicians are masters at making us fall into the trap of hidden assumptions. Imagine a radio commercial for a new drug in which someone says, "Baroxan lowered my blood pressure tremendously." If you think that sounds good, you have made a number of assumptions without being aware of them—in other words, hidden assumptions. Who says, for instance, that lowering blood pressure "tremendously" is a good thing (dead people have tremendously low blood pressure) or that the speaker's blood pressure was too high to begin with?

Magic, too, involves hidden assumptions. The trick in some magic acts is to make you assume that something happens, often by planting a false assumption in your mind. A magician might ask, "How did I move the ball from here to there?" while in reality he is using two balls. I won't knowingly put false assumptions into *your* mind in this book, but on occasion you and I (or you and your instructor) may unknowingly make different assumptions during a given discussion, a situation that unavoidably leads to confusion and misunderstanding. Therefore it is important that we carefully analyze our thinking and watch for the assumptions that we build into our models.

If the prediction of a hypothesis fails to agree with observations made to test the hypothesis, there are several ways to address the discrepancy. One way is to rerun the test to see if it is reproducible. If the test keeps producing the same result, it becomes necessary to revise the hypothesis, rethink the assumptions that went into it, or reexamine the original observations that led to the hypothesis.

#### Exercise 1.2 Dead music player

A battery-operated portable music player fails to play when it is turned on. Develop a hypothesis explaining why it fails to play, and then make a prediction that permits you to test your hypothesis. Describe two possible outcomes of the test and what you conclude from the outcomes. (*Think* before you peek at the answer below.)

**SOLUTION** There are many reasons the player might not turn on. Here is one example. Hypothesis: The batteries are dead. Prediction: If I replace the batteries with new ones, the player should work. Possible outcomes: (1) The player works once the new batteries are installed, which means the hypothesis is supported; (2) the player doesn't work after the new batteries are installed, which means the hypothesis is not supported and must be either modified or discarded.

**1.2** In Exercise 1.2, each of the conclusions drawn from the two possible outcomes contains a hidden assumption. What are the hidden assumptions?

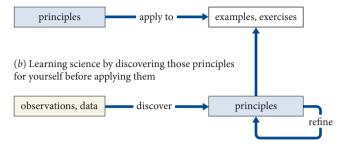
The development of a scientific hypothesis is often more complicated than suggested by Figures 1.1 and 1.2. Hypotheses do not always start with observations; some are developed from incomplete information, vague ideas, assumptions, or even complete guesses. The refining process also has its limits. Each refinement adds complexity, and at some point the complexity outweighs the benefit of the increased accuracy. Because we like to think that the universe has an underlying simplicity, it might be better to scrap the hypothesis and start anew.

Figure 1.2 gives an idea of the skills that are useful in doing science: interpreting observations, recognizing patterns, making and recognizing assumptions, thinking logically, developing models, and using models to make predictions. It should not come as any surprise to you that many of these skills are useful in just about any context. Learning physics allows you to sharpen these skills in a very rigorous way. So, whether you become a financial analyst, a doctor, an engineer, or a research scientist (to name just a few possibilities), there is a good reason to take physics.

Figure 1.1 also shows that doing science—and physics in particular-involves two types of reasoning: inductive, which is arguing from the specific to the general, and *deductive*, arguing from the general to the specific. The most creative part of doing physics involves inductive reasoning, and this fact sheds light on how you might want to learn physics. One way, which is neither very useful nor very satisfying, is for me to simply tell you all the general principles physicists presently agree on and then for you to apply those principles in examples and exercises (Figure 1.3a). This approach involves deductive reasoning only and robs you of the opportunity to learn the skill that is the most likely to benefit your career: discovering underlying patterns. Another way is for me to present you with data and observations and make you part of the discovery and refinement of the physics principles (Figure 1.3b). This approach

#### Figure 1.3

(a) Learning science by applying established principles



is more time-consuming, and sometimes you may wonder why I'm not just *telling* you the final outcome. The reason is that discovery and refinement are at the heart of doing physics!

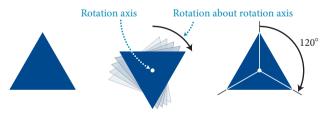
**1.3** After reading this section, reflect on your goals for this course. Write down what you would like to accomplish and why you would like to accomplish this. Once you have done that, turn to the final section of the *Principles* volume, "Solutions to checkpoints," and compare what you have written with what I wrote.

#### 1.2 Symmetry

One of the basic requirements of any law of the universe involves what physicists call *symmetry*, a concept often associated with order, beauty, and harmony. We can define **symmetry** as follows: An object exhibits symmetry when certain operations can be performed on it without changing its appearance. Consider the equilateral triangle in **Figure 1.4a**. If you close your eyes and someone rotates the triangle by 120° while you have your eyes closed, the triangle appears

#### Figure 1.4

(*a*) Rotational symmetry: Rotating an equilateral triangle by 120° doesn't change how it looks



(*b*) Reflection symmetry: Across each reflection axis (labeled R), two sides of the triangle are mirror images of each other

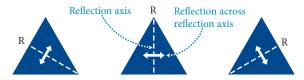
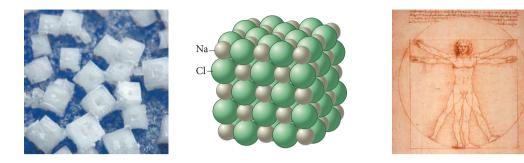


Figure 1.5 The symmetrical arrangement of atoms in a salt crystal gives these crystals their cubic shape.

(a) Micrograph of salt crystals

(*b*) Symmetrical arrangement of atoms in a salt crystal



the same when you open your eyes, and you can't tell that it has been rotated. The triangle is said to have *rotational symmetry*, one of several types of geometrical symmetry.

Another common type of geometrical symmetry, *reflection symmetry*, occurs when one half of an object is the mirror image of the other half. The equilateral triangle in Figure 1.4 possesses reflection symmetry about the three axes shown in Figure 1.4*b*. If you imagine folding the triangle in half over each axis, you can see that the two halves are identical. Reflection symmetry occurs all around us: in the arrangement of atoms in crystals (Figure 1.5*a* and *b*) and in the anatomy of most life forms (Figure 1.5*c*), to name just two examples.

The ideas of symmetry-that something appears unchanged under certain operations-apply not only to the shape of objects but also to the more abstract realm of physics. If there are things we can do to an experiment that leave the result of the experiment unchanged, then the phenomenon tested by the experiment is said to possess certain symmetries. Suppose we build an apparatus, carry out a certain measurement in a certain location, then move the apparatus to another location, repeat the measurement, and get the same result in both locations.\* By moving the apparatus to a new location (translating it) and obtaining the same result, we have shown that the observed phenomenon has translational symmetry. Any physical law that describes this phenomenon must therefore mathematically exhibit translational symmetry; that is, the mathematical expression of this law must be independent of the location.

Likewise, we expect any measurements we make with our apparatus to be the same at a later time as at an earlier time; that is, translation in time has no effect on the measurements. The laws describing the phenomenon we are studying must therefore mathematically exhibit symmetry under translation in time; in other words, the mathematical expression of these laws must be independent of time.

(*c*) Da Vinci's Vitruvian Man shows the reflection symmetry

of the human body

#### Exercise 1.3 Change is no change

**Figure 1.6** shows a snowflake. Does the snowflake have rotational symmetry? If yes, describe the ways in which the flake can be rotated without changing its appearance. Does it have reflection symmetry? If yes, describe the ways in which the flake can be split in two so that one half is the mirror image of the other.

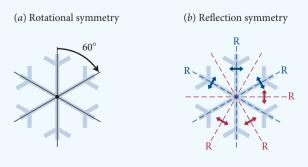
Figure 1.6 Exercise 1.3.



**SOLUTION** I can rotate the snowflake by 60° or a multiple of 60° (120°, 180°, 240°, 300°, and 360°) in the plane of the photograph without changing its appearance (**Figure 1.7***a*). It therefore has rotational symmetry.

I can also fold the flake in half along any of the three blue axes and along any of the three red axes in Figure 1.7*b*. The flake therefore has reflection symmetry about all six of these axes.

#### Figure 1.7



<sup>\*</sup>In moving our apparatus, we must take care to move any relevant external influences along with it. For example, if Earth's gravity is of importance, then moving the apparatus to a location in space far from Earth does not yield the same result.

A number of such symmetries have been identified, and the basic laws that govern the inner workings of the physical world must reflect these symmetries. Some of these symmetries are familiar to you, such as translational symmetry in space or time. Others, like electrical charge or parity symmetry, are unfamiliar and surprising and go beyond the scope of this course. Whereas symmetry has always implicitly been applied to the description of the universe, it plays an increasingly important role in physics: In a sense the quest of physics in the 21st century is the search for (and test of) symmetries because these symmetries are the most fundamental principles that all physical laws must obey.

**1.4** You always store your pencils in a cylindrical case. One day while traveling in the tropics, you discover that the cap, which you have placed back on the case day in, day out for years, doesn't fit over the case. What do you conclude?

#### 1.3 Matter and the universe

The goal of physics is to describe all that happens in the universe. Simply put, the universe is the totality of matter and energy combined with the space and time in which all events happen—everything that is directly or indirectly observable. To describe the universe, we use *concepts*, which are ideas or general notions used to analyze natural phenomena.\* To provide a quantitative description, these concepts must be expressed quantitatively, which requires defining a procedure for measuring them. Examples are the length or mass of an object, temperature, and time intervals. Such physical quantities are the cornerstones of physics. It is the accurate measurement of physical quantities that has led to the great discoveries of physics. Although many of the fundamental concepts we use in this book are familiar ones, quite a few are difficult to define in words, and we must often resort to defining these concepts in terms of the procedures used to measure them.

The fundamental physical quantity by which we map out the universe is **length**—a distance or an extent in space. The length of a straight or curved line is measured by comparing the length of the line with some standard length. In 1791, the French Academy of Sciences defined the standard unit for length, called the **meter** and abbreviated m, as one ten-millionth of the distance from the equator to the North Pole. For practical reasons, the standard was redefined in 1889 as the distance between two fine lines engraved on a bar of platinum-iridium alloy kept at the International Bureau of Weights and Measures near Paris. With the advent of lasers, however, it became possible to measure the speed of light with extraordinary accuracy, and so the meter was redefined in 1983 as the distance traveled

\*When an important concept is introduced in this book, the main word pertaining to the concept is printed in **boldface type**. All important concepts introduced in a chapter are listed at the end of the chapter, in the Chapter Glossary. by light in vacuum in a time interval of 1/299,792,458 of a second. This number is chosen so as to make the speed of light exactly 299,792,458 meters per second and yield a standard length for the meter that is very close to the length of the original platinum-iridium standard. This laser-based standard is final and will never need to be revised.

**1.5** Based on the early definition of the meter, one tenmillionth of the distance from the equator to the North Pole, what is Earth's radius?

Now that we have defined a standard for length, let us use this standard to discuss the structure and size scales of the universe. Because of the extraordinary range of size scales in the universe, we shall round off any values to the nearest power of ten. Such a value is called an order of magnitude. For example, any number between 0.3 and 3 has an order of magnitude of 1 because it is within a factor of 3 of 1; any number greater than 3 and equal to or less than 30 has an order of magnitude of 10. You determine the order of magnitude of any quantity by writing it in scientific notation and rounding the coefficient in front of the power of ten to 1 if it is equal to or less than 3 or to 10 if it is greater than 3.<sup>†</sup> For example, 3 minutes is 180 s, which can be written as  $1.8 \times 10^2$  s. The coefficient, 1.8, rounds to 1, and so the order of magnitude is  $1 \times 10^2$  s =  $10^2$  s. The quantity 680, to take another example, can be written as  $6.8 \times 10^2$ ; the coefficient 6.8 rounds to 10, and so the order of magnitude is  $10 \times 10^2 = 10^3$ . And Earth's circumference is 40,000,000 m, which can be written as  $4 \times 10^7$  m; the order of magnitude of this number is 10<sup>8</sup> m. You may think that using order-ofmagnitude approximations is not very scientific because of the lack of accuracy, but the ability to work effectively with orders of magnitude is a key skill not just in science but also in any other quantitative field of endeavor.

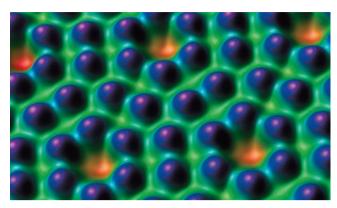
All ordinary matter in the universe is made up of basic building blocks called **atoms** (**Figure 1.8**). Nearly all the matter in an atom is contained in a dense central nucleus, which consists of **protons** and **neutrons**, two types of subatomic particles. A tenuous *cloud* of **electrons**, a third type of subatomic particle, surrounds this nucleus. Atoms are spherical and have a diameter of about  $10^{-10}$  m. Atomic nuclei are also spherical, with a diameter of about  $10^{-15}$  m, making atoms mostly empty space. Atoms attract one another when they are a small distance apart but resist being squeezed into one another. The arrangement of atoms in a material determines the properties of the material.

**Figure 1.9** shows the relative size of some representative objects in the universe. The figure reveals a lot about the organization of matter in the universe and serves as a visual model of the structure of the universe. Roughly

<sup>&</sup>lt;sup>†</sup>The reason we use 3 in order-of-magnitude rounding, and not 5 as in ordinary rounding, is that orders of magnitude are logarithmic, and on this logarithmic scale log 3 = 0.48 lies nearly halfway between log 1 = 0 and log 10 = 1.

**Figure 1.8** Scanning tunneling microscope image showing the individual atoms that make up a silicon surface. The size of each atom is about 1/50,000 the width of a human hair.

Figure 1.9 A survey of the size and structure of the universe.



speaking, there is clustering of matter from smaller to larger at four length scales. At the subatomic scale, most of the matter in an atom is compressed into the tiny atomic nucleus, a cluster of subatomic particles. Atoms, in turn, cluster to form the objects and materials that surround us, from viruses to plants, animals, and other everyday objects. The next level is the clustering of matter in stars, some of which, such as the Sun, are surrounded by planets like Earth. Stars, in turn cluster to form galaxies. As we shall discuss in Chapter 7, this clustering of matter reveals a great deal about the way different objects interact with one another.

#### **Exercise 1.4 Tiny universe**

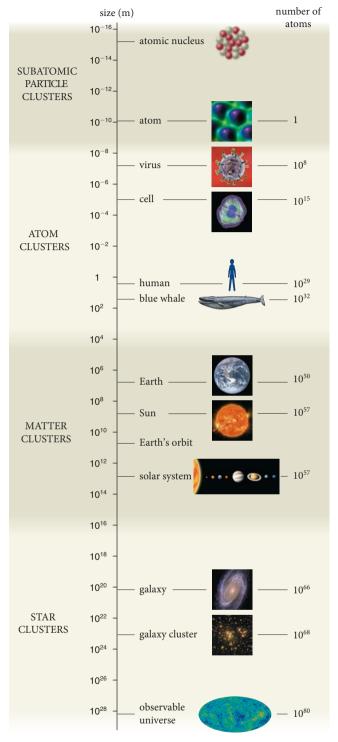
If all the matter in the observable universe were squeezed together as tightly as the matter in the nucleus of an atom, what order of magnitude would the diameter of the universe be?

**SOLUTION** From Figure 1.9 I see that there are about  $10^{80}$  atoms in the universe. I can arrange these atoms in a cube that has  $10^{27}$  atoms on one side because such a cube could accommodate  $10^{27} \times 10^{27} \times 10^{27} = 10^{81}$  atoms. Given that the diameter of a nucleus is about  $10^{-15}$  m, the length of a side of this cube would be

$$(10^{27} \text{ atoms})(10^{-15} \text{ m per atom}) = 10^{12} \text{ m},$$

which is a bit larger than the diameter of Earth's orbit around the Sun.

An alternative method for obtaining the answer is to realize that the matter in a single nucleus occupies a cubic volume of about  $(10^{-15} \text{ m})^3 = 10^{-45} \text{ m}^3$ . If all the matter in the universe were squeezed together just as tightly, it would occupy a volume of about  $10^{80}$  times the volume of an atomic nucleus, or  $10^{80} \times 10^{-45} \text{ m}^3 = 10^{35} \text{ m}^3$ . The side of a cube of this volume is equal to the cube root of  $10^{35} \text{ m}^3$ , or  $4.6 \times 10^{11} \text{ m}$ , which is the same order of magnitude as my first answer.



**1.6** Imagine magnifying each atom in an apple to the size of the apple. What would the diameter of the apple then be?

#### 1.4 Time and change

Profound and mysterious, time is perhaps the greatest enigma in physics. We all know what is meant by time, but it is difficult, if not impossible, to explain the idea in words. (Put the book down for a minute and try defining time in words before reading on.) One way to describe time is that it is the infinite continual progression of events in the past, present, and future, often experienced as a force that moves the world along. This definition is neither illuminating nor scientifically meaningful because it merely relates the concept of time to other, even less well-defined notions. Time is defined by the rhythm of life, by the passing of days, by the cycle of the seasons, by birth and death. However, even though many individual phenomena, such as the 24-hour cycle of the days, the cycle of seasons, and the swinging of a pendulum, are repetitive, the time we experience does not appear to be repetitive, and the current view is that time is a continuous succession of events.

The irreversible flow of time controls our lives, pushing us inexorably forward from the past to the future. Whereas we can freely choose our location and direction in all three dimensions in space, time flows in a single direction, dragging us forward with it. Time thus presents less symmetry than the three dimensions of space: Although opposite directions in space are equivalent, opposite directions in time are not equivalent. The "arrow of time" points only into the future, a direction we define as the one we have no memory of. Curiously, most of the laws of physics have no requirement that time has to flow in one direction only, and it is not until Chapter 19 that we can begin to understand why events in time are irreversible.

The arrow of time allows us to establish a *causal relation-ship* between events. For example, lightning causes thunder and so lightning has to occur *before* the thunder. This statement is true for all observers: No matter who is watching the storm and no matter where that storm is happening, every observer first sees a lightning bolt and only after that hears the thunder because an effect never precedes its cause. Indeed, the very organization of our thoughts depends on the **principle of causality**:

### Whenever an event A causes an event B, all observers see event A happening first.

Without this principle, it wouldn't be possible to develop any scientific understanding of how the world works. (No physics course to take!) The principle of causality also makes it possible to state a definition: **Time** is a physical quantity that allows us to determine the sequence in which related events occur.

To apply the principle of causality and sort out causes and effects, it is necessary to develop devices—clocks—for keeping track of time. All clocks operate on the same principle: They repeatedly return to the same state. The rotation of Earth about its axis can serve as a clock if we note the instant the Sun reaches its highest position in the sky on successive days. A swinging pendulum, which repeatedly returns to the same vertical position, can also serve as a clock. The time interval between two events can be determined by counting the number of pendulum swings between the events. The accuracy of time measurements can be increased by using a clock that has a large number of repetitions in a given time interval.

**1.7** (*a*) State a possible cause for the following events: (*i*) The light goes out in your room; (*ii*) you hear a loud, rumbling noise; (*iii*) a check you wrote at the bookstore bounces. (*b*) Could any of the causes you named have occurred after their associated event? (*c*) Describe how you feel when you experience an event but don't know what caused it—you hear a strange noise when camping, for instance, or an unexpected package is sitting on your doorstep.

The familiar standard unit for measuring time is the **second** (abbreviated s), originally defined as 1/86,400 of a day but currently more accurately defined as the duration of 9,192,631,770 periods of certain radiation emitted by cesium atoms. **Figure 1.10** gives an idea of the vast range of time scales in the universe.

The English physicist Isaac Newton stated, "Absolute, true, and mathematical time, of itself and from its own nature, flows equably without relation to anything external." In other words, the notion of past, present, and future is universal—"now" for you, wherever you are at this instant, is also "now" everywhere else in the universe. Although this notion of the universality of time, which is given the name **absolute time**, is intuitive, experiments described in Chapter 14 have shown this notion to be false. Still, for many experiments and for the material we discuss in most of this book, the notion of absolute time remains an excellent approximation.

Now that we have introduced space and time, we can use these concepts to study events. Throughout this book, we focus on **change**, the transition from one state to another. Examples of change are the melting of an ice cube, motion (a change in location), the expansion of a piece of metal, the flow of a liquid. As you will see, one might well call physics the study of the changes that surround us and convey the passage of time. What is most remarkable about all this is that we shall discover that underneath all the changes we'll look at, certain properties remain *unchanged*. These properties give rise to what are called *conservation laws*, the most fundamental and universal laws of physics.

There is a profound aesthetic appeal in knowing that symmetry and conservation are the cornerstones of the laws that govern the universe. It is reassuring to know that an elegant simplicity underlies the structure of the universe and the relationship between space and time. **1.8** A single chemical reaction takes about  $10^{-13}$  s. What order of magnitude is the number of sequential chemical reactions that could take place during a physics class?

**Figure 1.10** A survey of the time scales on which events in the universe take place.

time interval (s) 10-22 10-20 period of nuclear vibration INFERRED 10-18 10-16 period of visible light  $10^{-14}$ period of molecular vibration  $10^{-12}$ OBSERVABLE  $10^{-10}$ duration of one computer cycle  $10^{-8}$  $10^{-6}$  $10^{-4}$ duration of camera flash  $10^{-2}$ period of human heartbeat 1 10<sup>2</sup> HUMAN EXPERIENCE 10<sup>4</sup> NOV day 30 10<sup>6</sup> 1 vear 10<sup>8</sup> human life span 10<sup>10</sup> time to ancient civilizations 10<sup>12</sup> HISTORY time to age of dinosaurs 10<sup>14</sup> age of Earth 10<sup>16</sup> age of the universe 10<sup>18</sup>

#### **1.5 Representations**

Of all our senses, it is vision that most informs our mind. For this reason, expert problem solvers rarely start working on a problem without first making some sort of visual representation of the available information, and you should always do the same. Such a drawing helps you establish a clear mental image of the situation, interpret the problem in light of your own knowledge and experience, develop a qualitative understanding of the problem, and organize the information in a meaningful way. Without the drawing, you have to juggle all the information in your head. The drawing frees your brain to deal with the solution. As an example, consider the situation described in Checkpoint 1.9. First try solving the problem without a piece of paper and note the mental effort it takes; then make a sketch and work out the solution.

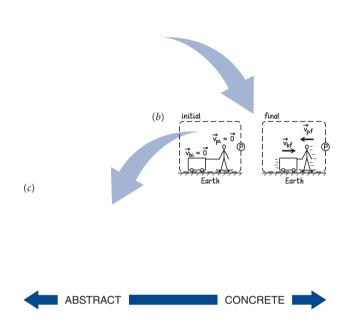
**1.9** You and your spouse are working out a seating arrangement around a circular table for dinner with John and Mary Jones, Mike and Sylvia Masters, and Bob and Cyndi Ahlers. Mike is not fond of the Ahlers, and Sylvia asked that she not be seated next to John. You would like to alternate men and women and avoid seating spouses next to each other. Determine an arrangement that satisfies all the constraints.

Were you able to solve the problem without a sketch? Doing so would be difficult because there is more information than most people can comfortably keep in mind at once. When you represent the information visually, however, it is relatively easy to solve this problem (see the solution in the back of this book). The drawing breaks the problem down into small steps and helps you articulate in your mind what you are trying to accomplish.

Visual representations are not helpful only for making seating arrangements. They work just as well for solving physics problems, although it may not immediately be clear *how* to represent the available information. For this reason we shall develop a number of different ways—pictures, sketches, diagrams, graphs—and a number of different context-specific procedures to represent information visually in our study of physics. As you will see, these visual representations are an integral part of getting a grip on a problem and developing a model (Figure 1.11).

As we discussed in Section 1.1, descriptions of the physical world always begin with simplified representations. When you solve physics problems, elaborate drawings clutter your mind with irrelevant information and prevent you from getting a clear view of the important features. One of the most basic skills in physics, therefore, is to decide what to leave out of your drawings. If you leave out essential features, the representation is useless, but if you put in too many details, it becomes impossible to analyze the situation. Sometimes it is necessary to begin with an oversimplification in order to develop a feel for a given situation. Once this initial understanding has been gained, however, **Figure 1.11** Multiple representations help you solve problems. (*a*) Many problems start with a verbal representation. (*b*) Turning the words into a sketch helps you to grasp the problem. (*c*) The sketch can then give meaning to a mathematical representation of the problem.

(a)



it becomes possible to construct less idealized models, with each successive model being a more realistic representation of the real-world situation.

In any drawings you make, therefore, you should treat everyday objects as simplifications that can be characterized by a minimum number of features or quantities. Some joke that physicists hold a grossly oversimplified view of the world, thinking in such terms as, for instance, "Consider a spherical cow...." The world around us, cows included, contains infinitely many details that may play a role in the grand scheme of things, but to get a grip on any problem it is important to begin by leaving out as much detail as possible. If the resulting model reproduces the main features of the real world, you know you have taken the essential attributes into account. As an example, **Figure 1.12** shows a progression from a photograph of a cow to an abstract rendition of it. To study the pattern on the cow's hide, you need the photograph. If you are interested in only the position of a certain cow in a certain pasture, however, a simple dot suffices (Figure 1.12*d*). By reducing the cow to a dot, you have discarded any information about its size and shape, but as you will see as you continue with your study of physics, this information is often not relevant.

#### **Exercise 1.5 Stretching a spring**

For a physics laboratory assignment, one end of a spring is attached to a horizontal rod so that the spring hangs vertically, and a ruler is hung vertically alongside the spring. The stretching properties of the spring are to be measured by attaching eight identical beads to the spring's free end. With no beads attached, the free end of the spring is at a ruler reading of 23.4 mm. With one bead attached, the end of the spring drops to 25.2 mm. When the second, third, and fourth beads are attached one at a time, the end drops to ruler readings of 26.5 mm, 29.1 mm, and 30.8 mm, respectively. Adding the fifth and sixth beads together moves the spring end to 34.3 mm, and adding the last two beads moves the end to 38.2 mm. (*a*) Make a pictorial representation of this setup. (b) Tabulate the data. (c) Plot the data on a graph, showing the ruler readings on the vertical axis and the numbers of beads on the horizontal axis. (d) Describe what can be inferred from the data.

**SOLUTION** (*a*) The important items that should appear in my drawing are the spring, the rod from which it is suspended, the ruler, and at least one bead. My drawing should also indicate how the ruler readings are obtained. I can illustrate this procedure with just one bead, as **Figure 1.13a** shows. By drawing the spring first with no beads attached and then with one bead attached and showing the two ruler readings, I've represented the general procedure of adding beads one (or two) at a time and paying attention to how each addition changes the position of the spring end.

(*b*) See Figure 1.13*b*.

(*c*) See Figure 1.13*c*.

(*d*) The relationship between the ruler readings and the numbers of suspended beads is linear. That is to say, each additional bead stretches the spring by about the same amount.

Figure 1.12 Increasingly simplified and abstract representations of a cow.

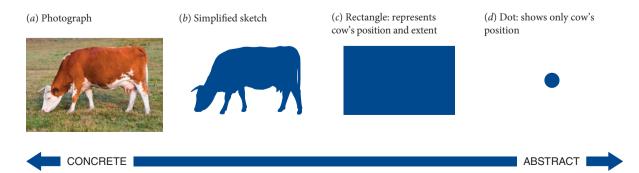
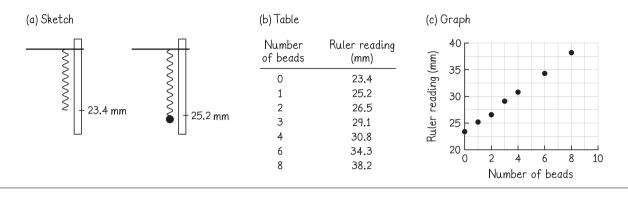


Figure 1.13



Exercise 1.5 demonstrates how various ways of representing information help you interpret data. The pictorial representation of Figure 1.13a helps you visualize how the measurements were taken. The table of Figure 1.13borganizes the data, and the graph of Figure 1.13c allows you to recognize the linear relationship between how far the spring stretches and the number of beads suspended from it—something that is not at all obvious from the text or the table.

Like the sketch, the graph is a simplified representation of the stretching of the spring. Each representation involves a loss of information and detail. The sketch is a simplified two-dimensional representation of a three-dimensional setup, and the graph shows only one piece of information for each measurement: the position of the bottom of the spring. All other information is left out in order to reveal one crucial point: How much the spring stretches is proportional to the number of beads suspended from it (something we look at in more detail in Chapter 9).

You may be beginning to wonder what the role of mathematics is in physics, given that we haven't used any thus far. One of the main roles of mathematics in physics is allowing us to express succinctly and unambiguously ideas that if expressed verbally would require a lot of words whose meaning may not be precise. Take this statement, for example:

The magnitude of the acceleration of an object is directly proportional to the magnitude of the vector sum of the forces exerted on the object and inversely proportional to the object's inertia. The direction of the acceleration is the same as the direction of the vector sum of the forces.

(Don't worry about the meaning of this statement for now; just note that it is quite a mouthful.) We can express this

statement mathematically as

$$\vec{a} = \frac{\sum \vec{F}}{m}.$$

These two expressions—one verbal, the other mathematical—are identical. Neither one is more accurate than the other. Right now the symbols in the equation may make no more sense to you than hieroglyphs. Once you understand their meaning, however, the mathematical expression can be parsed much more quickly than the verbal expression can. In this respect, mathematics plays the same role as visual representations: It relieves the brain of having to keep track of many words. Another important benefit of mathematical representation is that we can use the techniques of mathematics to manipulate the symbols and obtain new insights.

Without an understanding of the meaning of the concepts in any expression (acceleration and force in our example here), however, verbal expressions and mathematical ones are both meaningless, and so it is important to focus first on the meanings of concepts. As you will notice, this book has been designed to develop concepts first and to emphasize the visual representation of these concepts before moving on to a mathematical treatment (see the box "Organization of this book").

**1.10** Picture a long, straight corridor running east-west, with a water fountain located somewhere along it. Starting from the west end of the corridor, a woman walks a short distance east along the corridor and stops before reaching the water fountain. The distance from her to the fountain is twice the distance she has walked. She then continues walking east, passes the water fountain, and stops 60 m from her first stop. Now the distance from her to the fountain is twice the distance from her to the corridor. How long is the corridor?

#### **Organization of this book**

The material for this course is presented in two volumes. The first volume (the one you are reading now, *Principles*) is aimed at guiding you in developing a solid understanding of the principles of physics. The second volume, *Practice*, provides a broad variety of questions and problems that allow you to apply and sharpen your understanding of physics.

Each chapter in the *Principles* volume is divided into two parts: *Concepts* and *Quantitative Tools*. The *Concepts* part develops the conceptual framework for the subject of the chapter, concentrating on the underlying ideas and helping you develop a mental picture of the subject. The *Quantitative Tools* part develops the mathematical framework, building on the ideas developed in the *Concepts* part. Interspersed in the text of each chapter are several types of learning aids to guide you through the chapter:

1. Checkpoints ( ). These questions compel you to test yourself on how well you understand the material you just read. Do not skip them as you read the text. First of all, the answer to a checkpoint may be necessary to understand the text following that checkpoint. Second, and more important, I've put these checkpoints right in the text because working on them means learning the material. The answers to all the checkpoints are at the end of this volume.

- 2. Exercises and examples. The fully-worked-out exercises and examples help you develop and apply problemsolving strategies. It is generally a good idea to attempt to solve the problem by yourself before reading the solution. More information on general approaches to problem solving is given in Section 1.8.
- **3. Procedures.** Approaches for analyzing specific situations are given in separate, highlighted boxes.
- **4. Self-quiz.** The Self-quiz in each *Principles* chapter, which always comes at the end of the conceptual part of the chapter, allows you to assess your understanding of the concepts before you move on to the quantitative treatment. Complete each Self-quiz before working on the quantitative part of the chapter, even if you are already familiar with most of the material covered. Before tackling the quantitative material, be sure to resolve any difficulties you might have in answering a Self-quiz question, either by rereading the material in the conceptual part of the chapter or by consulting your instructor.
- **5. Glossary**. At the end of each chapter is a list defining the important concepts in the chapter (which are the terms printed in bold).